

Collider and CMB Complementarity: An EFT Approach

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PHYSICAL
RESEARCH
LABORATORY

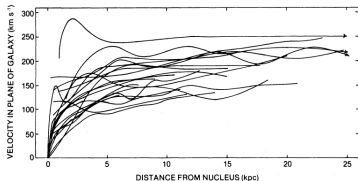
Vikram Discussions on Neutrino Astrophysics, 2025
PRL, Ahmedabad

Based on *arXiv:2408.14548, JHEP01(2025)074*
D. Borah, N. Das, S. Jahedi, and B. Thacker

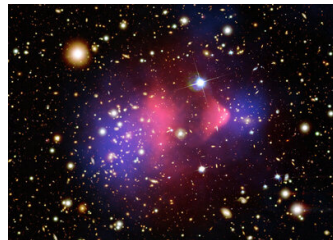
Outline

- New Physics and Where to find it?
- Effective Operators
- Dark Matter Phenomenology
- Cosmological Signatures
- Results and Conclusion

Dark Matter Evidences and Scale



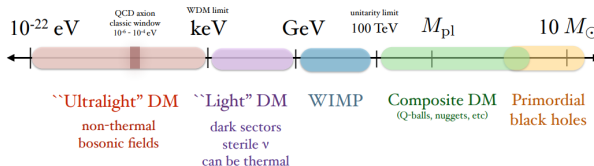
Vera C Rubin et al., *Astrophysical Journal*



The Bullet Cluster

Mass scale of dark matter

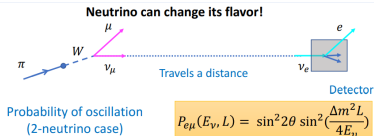
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Neutrino Physics

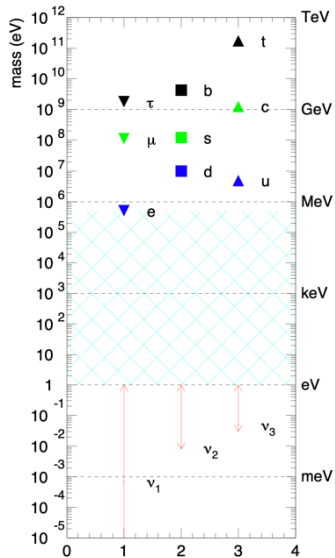
In the Standard Model:

- Three flavors of neutrinos
- Neutrinos are massless
- Neutrinos interact only weakly



Beyond the Standard Model:

- Neutrino oscillations → neutrinos have mass
- Nature: Dirac or Majorana?



Setup

- Dark Matter (χ): A Dirac fermion; $\mathcal{G}_{\text{DM}} = \mathbb{Z}_2$
- Right-Handed Neutrinos (ν_R): 3
- EFT Operators:

$$\mathcal{O}_{\text{DM}}\mathcal{O}_{\text{SM}}, \quad \mathcal{O}_{\text{DM}}\mathcal{O}_{\nu_R}, \quad \mathcal{O}_{\text{SM}-\nu_R}$$

- Symmetry Requirements:
 - Operators are invariant under $\mathcal{G}_{\text{SM}} = SU(3)_C \times SU(2)_L \times U(1)_Y$
 - Global lepton number $(\Delta L)_{\text{global}} = 0$ is preserved, preventing Majorana mass terms and ensuring the Dirac nature of ν_R .

Effective Operators

Dimension-6 Operators

The operators are of the form:

$$\frac{c_{\chi l}}{\Lambda^2} \mathcal{O}_{\chi l}, \quad \frac{c_{\chi\nu}}{\Lambda^2} \mathcal{O}_{\chi\nu}, \quad \frac{c_{\text{SM}\nu_R}}{\Lambda^2} \mathcal{O}_{\text{SM}-\nu_R}$$

where c_{ij} are Wilson coefficients and Λ is the scale of new physics.

Class	Name	\mathcal{O}_{DM}	$\mathcal{O}_{\text{SM}}, \mathcal{O}_{\nu_R}$
DMEFT	$\mathcal{O}_{\chi l}$	$(\bar{\chi}\gamma^\mu\chi), (\bar{\chi}\gamma^\mu\gamma^5\chi)$	$(\bar{L}\gamma_\mu L), (\bar{e}_R\gamma_\mu e_R)$
ν DMEFT	$\mathcal{O}_{\chi\nu}$	$(\bar{\chi}\gamma^\mu\chi), (\bar{\chi}\gamma^\mu\gamma^5\chi)$	$(\bar{\nu}_R\gamma_\mu\nu_R)$

Table: Operators involving DM, SM, and ν_R leading to interactions relevant for leptophilic DM phenomenology

Operators for SM- ν_R Interactions

No.	Name	Operator	No.	Name	Operator
1	$\mathcal{O}_{L\nu H}$	$(\bar{L}\nu_R)\tilde{H}(H^\dagger H) + \text{h.c.}$	9	$\mathcal{O}_{d\nu}$	$(\bar{d}_R\gamma^\mu d_R)(\bar{\nu}_R\gamma_\mu\nu_R)$
2	$\mathcal{O}_{H\nu}$	$(\bar{\nu}_R\gamma^\mu\nu_R)(H^\dagger i\overleftrightarrow{D}_\mu H)$	10	$\mathcal{O}_{du\nu e}$	$(\bar{d}_R\gamma^\mu u_R)(\bar{\nu}_R\gamma_\mu e_R) + \text{h.c.}$
3	$\mathcal{O}_{H\nu e}$	$(\bar{\nu}_R\gamma^\mu e_R)(H^\dagger iD_\mu H) + \text{h.c.}$	11	$\mathcal{O}_{L\nu}$	$(\bar{L}\gamma^\mu L)(\bar{\nu}_R\gamma_\mu\nu_R)$
4	$\mathcal{O}_{\nu B}$	$(\bar{L}\sigma_{\mu\nu}\nu_R)\tilde{H}B^{\mu\nu}$	12	$\mathcal{O}_{Q\nu}$	$(\bar{Q}\gamma^\mu Q)(\bar{\nu}_R\gamma_\mu\nu_R)$
5	$\mathcal{O}_{\nu W}$	$(\bar{L}\sigma_{\mu\nu}\nu_R)\tau^I\tilde{H}W^{I\mu\nu}$	13	$\mathcal{O}_{L\nu L e}$	$(\bar{L}^i\nu_R)\epsilon_{ij}(\bar{L}^j e_R) + \text{h.c.}$
6	$\mathcal{O}_{\nu\nu}$	$(\bar{\nu}_R\gamma^\mu\nu_R)(\bar{\nu}_R\gamma_\mu\nu_R)$	14	$\mathcal{O}_{L\nu Q d}$	$(\bar{L}^i\nu_R)\epsilon_{ij}(\bar{Q}^j d_R) + \text{h.c.}$
7	$\mathcal{O}_{e\nu}$	$(\bar{e}_R\gamma^\mu e_R)(\bar{\nu}_R\gamma_\mu\nu_R)$	15	$\mathcal{O}_{L d Q\nu}$	$(\bar{L}^i d_R)\epsilon_{ij}(\bar{Q}^j\nu_R) + \text{h.c.}$
8	$\mathcal{O}_{u\nu}$	$(\bar{u}_R\gamma^\mu u_R)(\bar{\nu}_R\gamma_\mu\nu_R)$	16	$\mathcal{O}_{Qu\nu L}$	$(\bar{Q}^i u_R)(\bar{\nu}_R L^i) + \text{h.c.}$

Table: Dimension-6 ν SMEFT operators. $\Delta L = 0, \Delta B = 0$

DM Relic Abundance and Freeze-Out

- **Thermal Equilibrium and Freeze-Out**
- **Free Parameters:** DM mass (m_χ) and cutoff scale (Λ), with Wilson coefficients $\mathcal{O}(1)$.
- **Boltzmann Equation:**

$$\frac{dY_\chi}{dx} = -1.32 \sqrt{g_*} M_{\text{pl}} \frac{m_\chi}{x^2} \langle \sigma v \rangle_{2\text{DM} \rightarrow 2\text{SM}} (Y_\chi^2 - Y_{\text{eq}}^2)$$

- **Relic Abundance:**

$$\Omega_\chi h^2 \propto 1/\langle \sigma v \rangle \propto \Lambda^4/m_\chi^2$$

- **Effects of New Operators:**
 - New annihilation channels delay freeze-out, reducing relic abundance.
 - $\mathcal{O}_{\chi\nu}$: parameter space shifted upward: increase in annihilation channels.

Relic Allowed Parameter Space

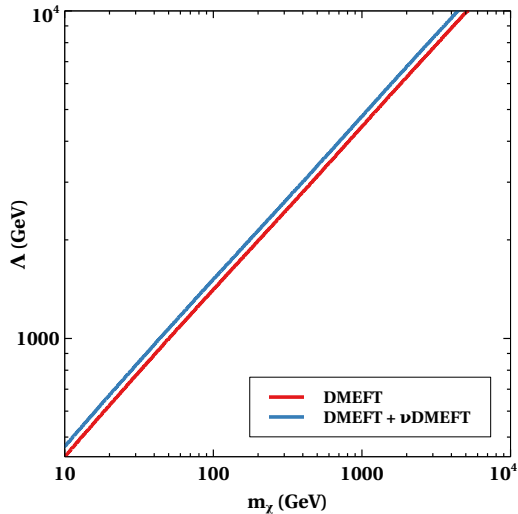


Figure: Relic allowed parameter space in $\Lambda - m_\chi$ plane

Direct Detection

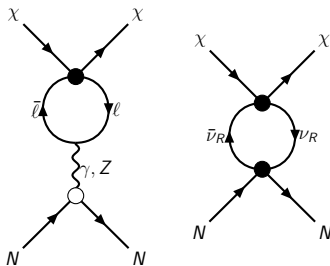


Figure: 1-loop Feynman diagrams from DM-nucleon scattering

The SI DM-nucleon cross-section for the $\mathcal{O}_{\chi\ell}$:

$$\sigma_{\chi N} \approx \frac{\mu_p^2}{9\pi A^2} \left(\frac{\alpha_{\text{em}} Z}{\pi \Lambda^2} \right)^2 \left[\ln \left(\frac{\Lambda^2}{\Lambda_{\text{DD}}^2} \right) \right]^2$$

Direct Detection

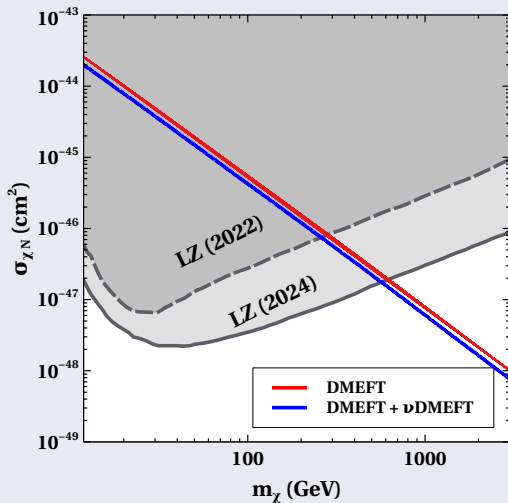


Figure: SI DM-nucleon cross-section for relic density allowed parameter space.

Indirect Search

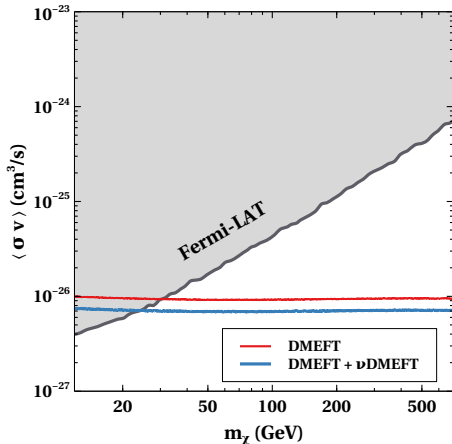


Figure: The shaded region is excluded from combined data analysis of MAGIC Cherenkov telescopes and Fermi-LAT experiment in $\langle \sigma v \rangle_{\chi\chi \rightarrow \tau^+\tau^-}$ versus m_χ plane.

Collider Phenomenology

Collider Signatures

- **Mono-X** + \cancel{E} , where X refers to γ , H, Jet, and Z.
- **Multi-lepton** / **Multi-Jet** channel: HDSP.

Hadron Colliders

- $\sqrt{\hat{s}}$ unknown. EFT validation questionable for $\sqrt{s} < \Lambda$.
- QCD background.
- Unpolarized beams

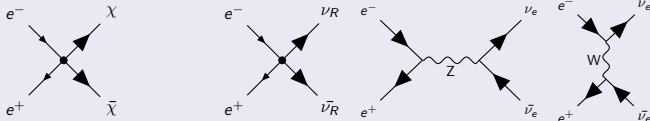
Lepton Colliders

- Known \sqrt{s} . EFT applicable for $\sqrt{s} < \Lambda$.
- Less QCD background.
- Initial beam polarization.

Tools

FeynRules (LanHEP), MadGraph, Pythia, Delphes, micrOMEGAs

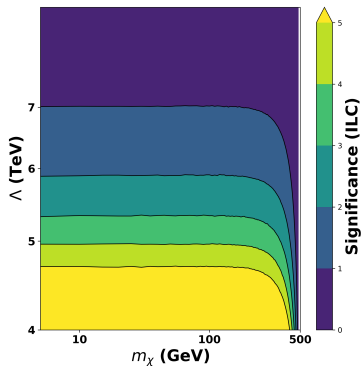
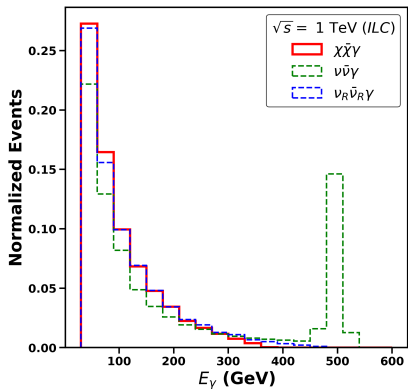
Collider Search: ILC



Beam Polarization	Cross-section (fb)		
	$\chi\bar{\chi}\gamma$ (S)	$\nu\bar{\nu}\gamma$ (B1)	$\nu_R\bar{\nu}_R\gamma$ (B2)
$\{0\%, 0\%\}$	12.34	2450	19.65
$\{+20\%, +80\%\}$	10.38	630	16.50
$\{-20\%, +80\%\}$	14.30	463	22.79
$\{+20\%, -80\%\}$	14.33	5218	22.92
$\{-20\%, -80\%\}$	10.35	3476	16.54

Table: $\sqrt{s} = 1$ TeV and relic satisfied EFT parameters, $m_\chi = 260$ GeV and $\Lambda = 2450$ GeV

Distributions and Significance



Left: E_γ distribution)
Right: Signal Significance

Cosmological Signatures

Key Points

- Light Dirac neutrinos: Contributing to ΔN_{eff}
- Radiation energy density:

$$\rho_r = \rho_\gamma + \rho_{\nu_L} + \rho_{\nu_R} = \left(1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{\text{eff}} \right) \rho_\gamma.$$

- SM prediction: $N_{\text{eff}}^{\text{SM}} = 3.045$. Deviations signal BSM physics.

Current and Future Bounds

- **Planck 2018:** $\Delta N_{\text{eff}} \lesssim 0.285$ at 2σ CL.
- **DESI 2024:** $\Delta N_{\text{eff}} \lesssim 0.4$ at 2σ CL.
- **CMB-S4 (Future):** $\Delta N_{\text{eff}} \sim 0.06$.
- **CMB-HD (Future):** $\Delta N_{\text{eff}} \sim 0.014$ (1σ).

Thermal Decoupling of ν_R

ΔN_{eff}

$$\Delta N_{\text{eff}} = 0.047 \times 3 \left(\frac{106.75}{g_{*\rho}(T_{\text{dec}})} \right)^{4/3},$$

$g_{*\rho}(T_{\text{dec}})$ is the relativistic DOF at ν_R decoupling.

Case	Interactions ($\bar{f}_{L,R}\gamma^\mu\psi_{L,R}$ type only)	Planck 2018 Limit on Λ	DESI 2024 Limit on Λ
1	All fermions	$\Lambda \lesssim 11.6$ TeV	$\Lambda \lesssim 7.8$ TeV
2	leptons	$\Lambda \lesssim 10$ TeV	$\Lambda \lesssim 6.7$ TeV
3	SM neutrinos	$\Lambda \lesssim 8$ TeV	$\Lambda \lesssim 5.2$ TeV
4	quarks	$\Lambda \lesssim 9.8$ TeV	$\Lambda \lesssim 6.5$ TeV
5	3rd Gen quarks	$\Lambda \lesssim 0.6$ TeV	$\Lambda \lesssim 0.001$ TeV
6	Dark matter (χ)	m_χ dependent	m_χ dependent

Table: ν_R - SM interactions and **disallowed values of Λ**

Constraints on ν_R Interactions

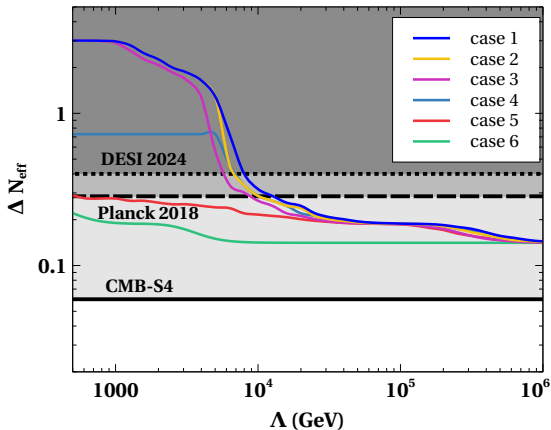
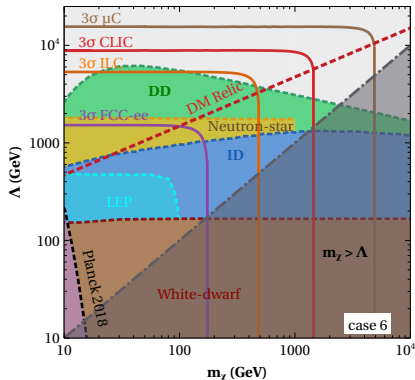
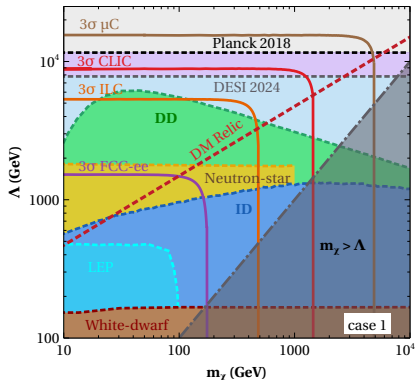


Figure: Variation of ΔN_{eff} with Λ

Summary Plots



Left: Case 1 (All fermions) **Right:** Case 6 (DM (χ))
 $\Lambda - m_\chi$ plane

Conclusion

- We have Studied **leptophilic DM** with light Dirac neutrinos using an **EFT approach**.
- ν DMEFT / ν SMEFT features **enhanced N_{eff}** measurable by future CMB experiments.
- BSM Discovery prospects explored at future lepton colliders (μ C, **CLIC, ILC, FCC-ee**) and CMB experiments (**Planck, DESI**).
- Collider sensitivities on Λ (\sim a few tens of TeV) and m_χ (\sim a few TeV).
- **Future CMB and Collider experiments can provide complementary searches for BSM Physics through EFT.**

Thank You!

Backup: ΔN_{eff} for ν_R Interactions

- **Interaction Rate:**

$$\Gamma = \sum_f \frac{n_f^{\text{eq}} n_{\bar{f}}^{\text{eq}}}{n_{\nu_R}^{\text{eq}}} \langle \sigma v \rangle_{f\bar{f} \rightarrow \nu_R \bar{\nu}_R}$$

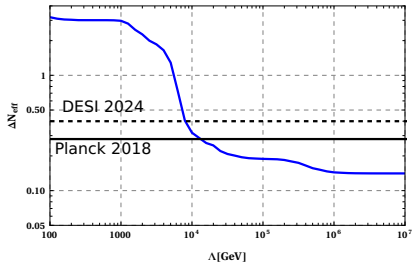
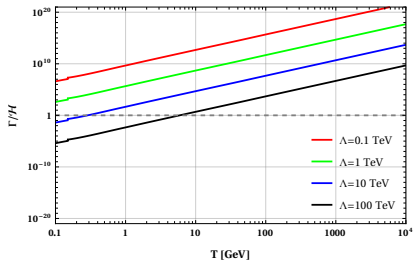
where $f(\bar{f})$: SM fermions and DM

- **Decoupling Temperature:**

$$\Gamma(T_{\text{dec}}) \sim \mathcal{H}(T_{\text{dec}})$$

ν_R decouples when interaction rate falls below Hubble expansion rate.

Backup: Results and Figures



● Results:

- **Left Panel:** Γ/\mathcal{H} vs T for various Λ
- **Right Panel:** ΔN_{eff} vs Λ
- Planck 2018 excludes $\Lambda \lesssim 11.6$ TeV for $\Delta N_{\text{eff}} > 0.28$
- DESI 2024 excludes $\Lambda \lesssim 7.8$ TeV

Back Up: Collider Details

Collider Specifications

Maximum reach of various lepton colliders, including \sqrt{s} and beam polarization.

Colliders	$(\sqrt{s}, \mathcal{L}_{\text{int}})$	Beam Polarization $\{P_{e^+}, P_{e^-}\}$
FCC-ee	$(365 \text{ GeV}, 340 \text{ fb}^{-1})$	—
ILC	$(1 \text{ TeV}, 8 \text{ ab}^{-1})$	$\{\pm 20\%, \pm 80\%\}$
CLIC	$(3 \text{ TeV}, 5 \text{ ab}^{-1})$	$\{\pm 0\%, \pm 80\%\}$
μC	$(10 \text{ TeV}, 10 \text{ ab}^{-1})$	—